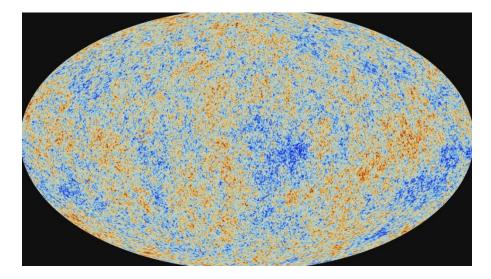
# ADVANCED COSMOLOGY

Màster: ASTROFÍSICA FÍSICA DE PARTÍCULES I COSMOLOGIA Universitat de Barcelona Jordi Miralda and Cristiano Germani



## PURPOSE

Introduce modern cosmology at an advanced enough level to prepare you for research.

- 1. Understand the basic elements of the expanding Universe, the Big Bang, and the Friedman-Robertson-Walker space-time metric.
- 2. Observational basis and theoretical framework of large-scale structure, dark matter and dark energy.
- 3. Origin of the Cosmic Microwave Background and the light element abundances.
- 4. Early Universe physics, phase transitions in cosmology.
- 5. Problems that lead to the inflationary paradigm, main features of cosmic inflation and origin of fluctuations.

## HOURS

The course is delivered over 13 weeks, 4 days per week, in lectures of one hour. First part (J. Miralda): Sep 13th – Oct 26th. Exam: Oct. 27th. Second part (C. Germani) starts Nov 2nd, exam Dec. 21st.

### **EVALUATION**

50 % by written exams (two, one each part) 50 % by assigned exercises

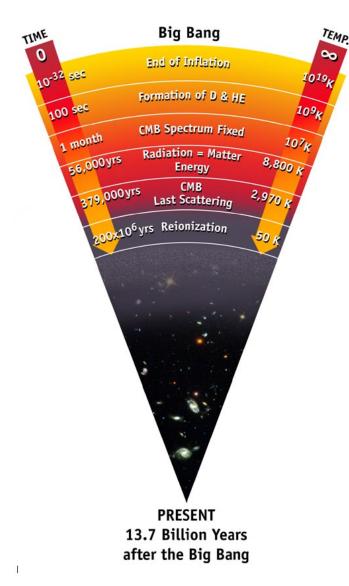
## **RECOMMENDED BACKGROUND**

Essential background in basic physics: Classical Mechanics, Special Relativity Electromagnetism, Thermodynamics, Statistical Mechanics

Some working knowledge at undergraduate level of:

Astrophysics and Cosmology General Relativity Quantum Mechanics

## FIRST PART



#### **1. The expansion of the Universe**

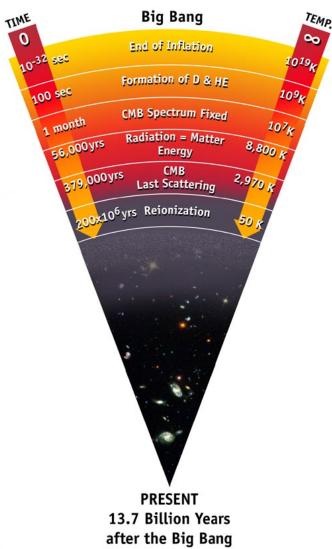
Spacetime geometry, expansion, Friedman equations, radiation, matter and cosmological constant

## 2. Observational cosmology and cosmic budget

Ages, cosmic abundances, baryon content, dark matter, dark energy, standard Lambda CDM model

**3. Cosmic microwave background** Fluctuations and power spectrum, dipole anisotropies, acoustic peaks, large-scale structure

## **SECOND PART**



# 4. Thermal history, nucleosynthesis and recombination Neutrino decoupling, neutron freeze-out, deuterium bottleneck, helium, recombination

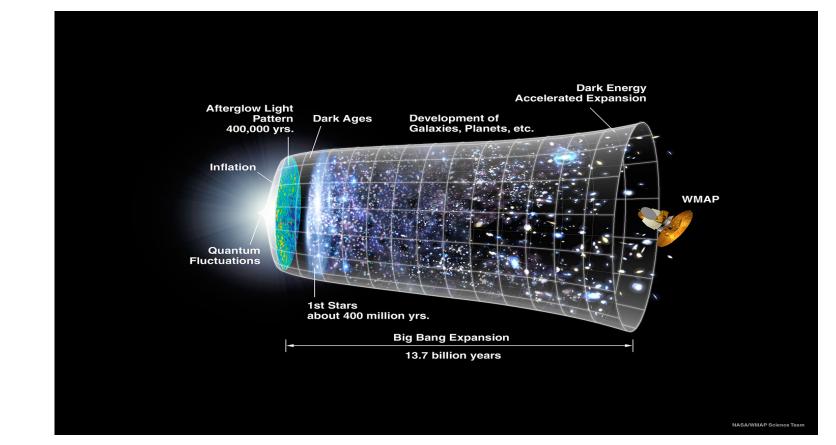
### 5. The very early universe

Standard model of particle physics, phase transitions, baryogenesis and leptogenesis, dark matter candidates.

#### 6. Inflation

Flatness, horizon problems, quintessence, slow roll inflation, thermalization.

## **BIBLIOGRAPHY**



V. Mukhanov, Physical foundations of cosmology, CUP, 2005
J.A. Peacock, Cosmological physics, CUP, 1990
B. Ryden, Introduction to cosmology, Addison Wesley, 2003